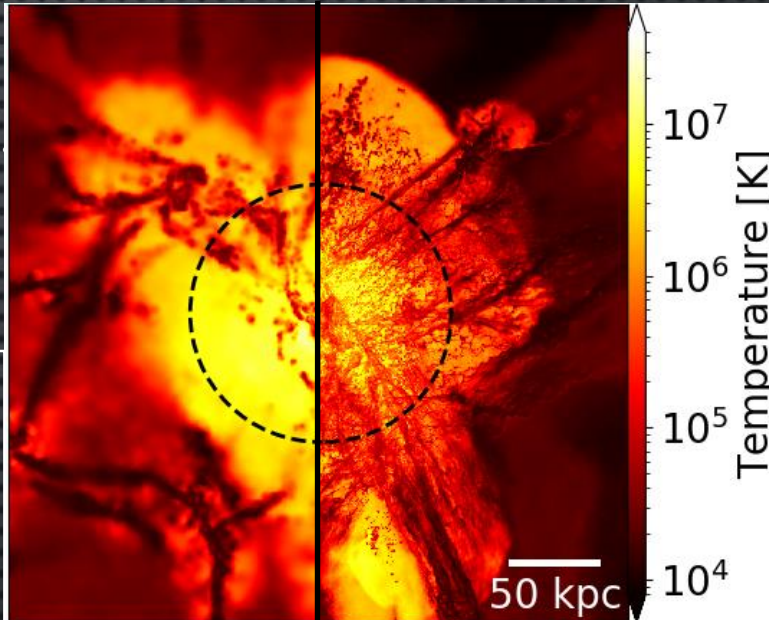
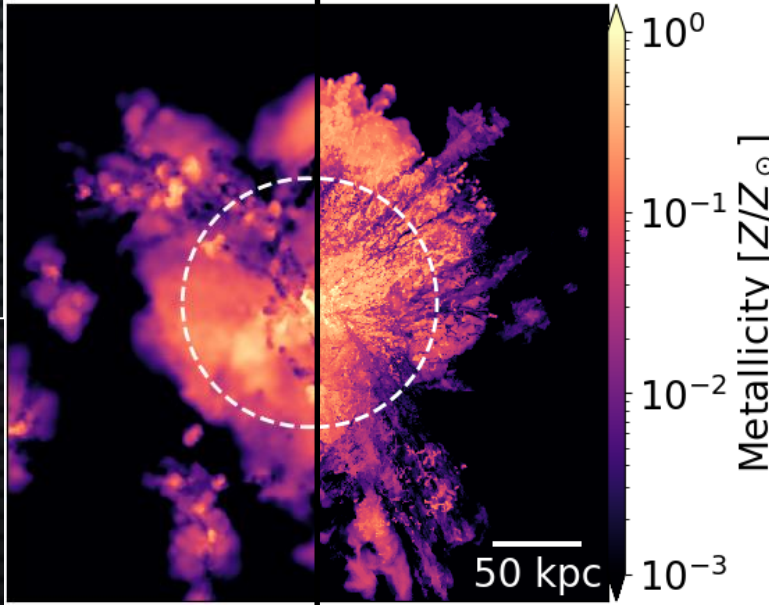


Recent observations have detected a multiphase circumgalactic medium, even in massive galaxies with significant shock heating (e.g. Rudie et al. 2012, Rubin et al. 2015). Hydrodynamical simulations have found that gas flows onto haloes take the form of cold filamentary accretion coexisting with a hot halo (e.g. Kereš et al. 2005, Dekel et al. 2009). However in typical simulations, resolution is poor in halo outskirts where this accretion occurs.



We have developed a novel shock refinement scheme within AREPO (Springel 2010), where resolution is increased on-the-fly near to shocks. For a  $10^{12} M_{\odot}$  halo at  $z=6$ , this particularly targets resolution at the accretion shocks (in yellow-white in the top-left figure), which are much sharper in the shock refined run (right) compared to the base run (left). Cool, metal-poor primordial filaments are much more prevalent within the hot halo when our scheme is active (shown in the figures on the right), leading to a much more multiphase CGM closer to observations.



# Resolving shocks & the multiphase CGM

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